

# Relaxation rates and phase transitions

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arXiv:1603.05950

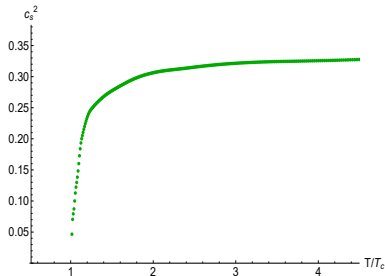
arXiv:1512.06871

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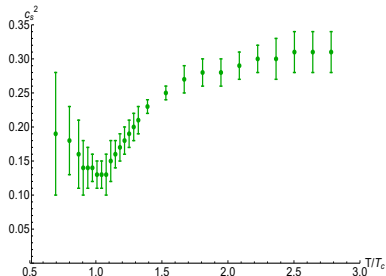


# Phase structure at strong coupling

- Systems at strong coupling exhibit various phase structures
- Pure gluon system  $\rightarrow$  1<sup>st</sup> order phase transition (left)
- Gluons + quarks  $\rightarrow$  smooth crossover (right)



G. Boyd *et al.* Nucl. Phys. B **469**,  
419 (1996)



S. Borsanyi *et al.* JHEP **1009**, 073  
(2010)

- Real time dynamics is not easy reachable with lattice methods
- Use other methods to model strongly coupled phase transitions
- Compute the spectrum of linearized perturbations
- Compute transport coefficients and non-hydrodynamic modes
- Check linear stability

## Method:

Use string theory based methods to formulate models at strong coupling!

- Does spinodal instability appear for a system with a 1<sup>st</sup> order phase transition?
- Does dynamical instability has to be accompanied by a thermodynamical instability?
- How non-hydrodynamic degrees of freedom behave in the critical region?
- Do diffusive modes appear?

## Method:

Use string theory based methods to formulate models at strong coupling!

# Holographic setup

- Add a source for an operator  $O_\phi$  in a  $\text{CFT}_{3+1}$

$$\mathcal{L} = \mathcal{L}_{\text{CFT}} + \Lambda^{4-\Delta} O_\phi$$

- Dual is a gravity-scalar theory in  $d = 5$

$$S = \frac{1}{2\kappa_5^2} \int_{\mathcal{M}} d^5x \sqrt{-g} \left[ R - \frac{1}{2} (\partial\phi)^2 - V(\phi) \right]$$

with the potential

$$V(\phi) = -12(1 + a\phi^2)^{1/4} \cosh(\gamma\phi) + b_2\phi^2 + b_4\phi^4 + b_6\phi^6$$

- In IHQCD family  $a \neq 0$  and  $T = 0$  geometry is *confining*

U. Gursoy, *et.al.* JHEP **0905**, 033 (2009)

S. S. Gubser, A. Nellore, Phys. Rev. D **78** (2008) 086007

# Phase transitions in holography

- Finite  $T$  states correspond to various black hole solutions in the dual spacetime
- Phase structure is determined by the choice of  $a, \gamma$  and  $b_2, b_4, b_6$ , coefficients of  $V(\phi)$
- It is possible to tune those parameters to mimic
  - crossover e.g. QCD
  - 1<sup>st</sup> order phase transition e.g. pure gluon systems
  - 2<sup>nd</sup> order phase transition

U. Gursoy, *et.al.* JHEP **0905**, 033 (2009)

S. S. Gubser, A. Nellore, Phys. Rev. D **78** (2008) 086007

# Linear response and Quasinormal modes

- Perturb the system  $\mathcal{L} = \mathcal{L}_0 + h_{ij}\delta^3(x)\delta(t)T^{ij}(x)$  the response is the *retarded Green's* function

$$G_R(\omega, k) \propto i \int dt d^3x \theta(t) e^{ikx - i\omega t} \langle [T_{ij}(x, t), T_{kl}(0)] \rangle$$

- *Quasinormal modes*, i.e., solutions of linearized fluctuation equations correspond to poles of holographic retarded Green's functions. In general

$$\omega_n(k) = \Omega_n(k) - i\Gamma_n(k)$$

where  $n = 1, 2, 3, \dots$   $\Omega_n(k)$ —oscillation frequency,  
 $\Gamma_n(k)$ —damping rate. Stable modes have  $\Gamma_n(k) > 0$ .

P. K. Kovtun, A. O. Starinets, Phys. Rev. D **72**, 086009 (2005)

# Linear response and Quasinormal modes

- Hydrodynamic mode is defined by

$$\lim_{k \rightarrow 0} \omega_H(k) = 0$$

- The sound mode

$$\omega(k) = \pm c_s k - \frac{i}{2T} \left( \frac{4}{3} \frac{\eta}{s} + \frac{\zeta}{s} \right) k^2 + O(k^3)$$

$\eta$ —shear viscosity,  $\zeta$ —bulk viscosity,  $s$ —entropy density,  
 $c_s$ —speed of sound,  $T$ —temperature

- In holographic models also *non-hydrodynamic* modes are present

P. K. Kovtun, A. O. Starinets, Phys. Rev. D **72**, 086009 (2005)  
M. P. Heller, R. A. Janik, P. Witaszczyk,  
Phys. Rev. Lett. **110**, no. 21, 211602 (2013)



# The first example $\rightarrow$ IHQCD

- Improved Holographic QCD

$$V(\phi) = -12(1 + a\phi^2)^{1/4} \cosh(\gamma\phi) + b_2\phi^2 + b_4\phi^4 + b_6\phi^6$$

with  $a = 1$ ,  $\gamma = \sqrt{2/3}$ ,  $b_2 = 6.25$ ,  $b_4 = b_6 = 0$

- Holographic model motivated by gluon dynamics
- In the deep IR the asymptotic form of the potential

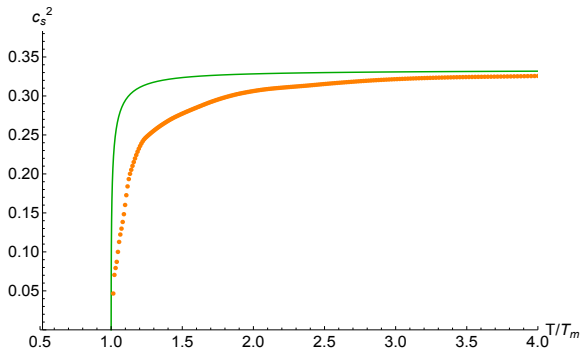
$$V(\phi) \sim -12\sqrt{\phi}e^{\gamma\phi}$$

determines a confining vacuum with asymptotic glueball spectrum  $m_n^2 \sim n$

U. Gursoy, *et.al.* JHEP **0905**, 033 (2009)

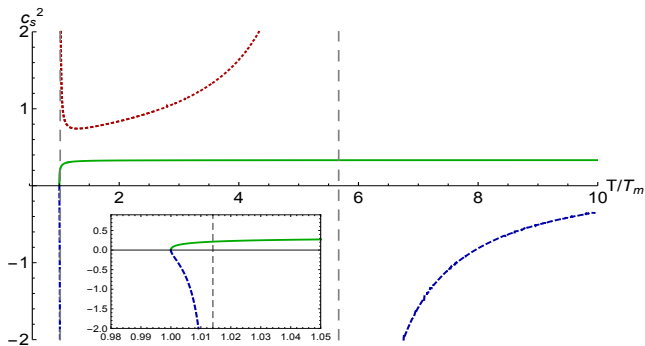
# The first example $\rightarrow$ IHQCD

- Transition between black hole and horizon-less geometry  
S. W. Hawking, D. N. Page, Commun. Math. Phys. **87**, 577 (1983)
- Holographic 1<sup>st</sup> order phase transition



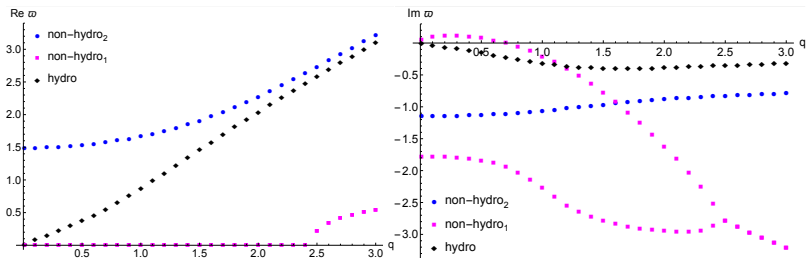
G. Boyd *et.al.* Nucl. Phys. B **469**, 419 (1996)

# Full holographic scan



- Below  $T_m$  no black hole solution exists
- Various lines represent different black hole phases with different properties

# Dynamical instability



- Quasinormal modes at  $T = 1.027 T_m$
- System displays dynamical instability despite thermodynamical stability!

- When  $c_s^2 < 0$  we have purely damped hydro-modes

$$\omega \approx \pm i|c_s|k - \frac{i}{2T} \left( \frac{4}{3} \frac{\eta}{s} + \frac{\zeta}{s} \right) k^2$$

so for small enough  $k$  we have  $\text{Im } \omega > 0$

- For a finite range of momenta this mode is present
- This appears for systems with a 1<sup>st</sup> order phase transition; *spinodal* instability
- This phenomenon occurs e.g. in nuclear matter

P. Chomaz, M. Colonna, J. Randrup, Phys. Rept. **389**, 263 (2004)

# The second example $\rightarrow$ first order phase transition

- The following choice

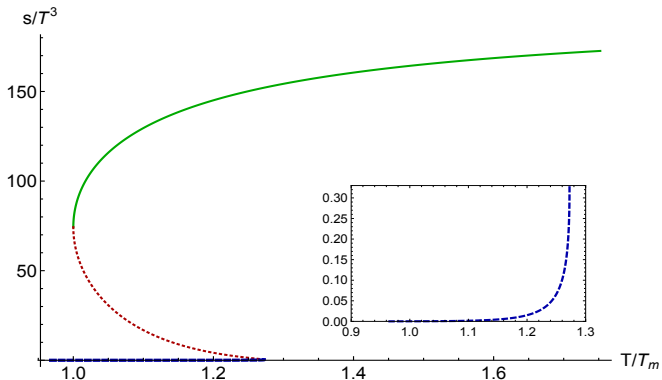
$$V(\phi) = -12(1 + a\phi^2)^{1/4} \cosh(\gamma\phi) + b_2\phi^2 + b_4\phi^4 + b_6\phi^6$$

with  $a = 0$ ,  $\gamma = \sqrt{7/12}$ ,  $b_2 = 2.5$ ,  $b_4 = b_6 = 0$

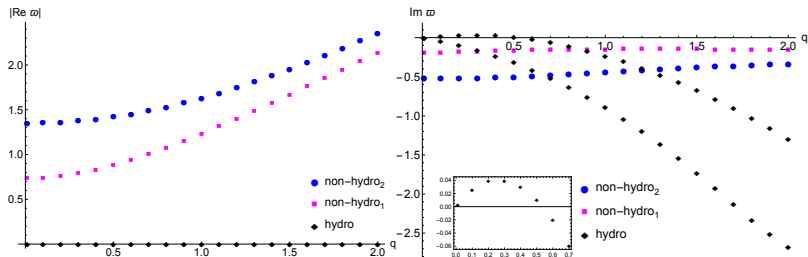
- Transition between two different black hole solutions
- Other example of holographic 1<sup>st</sup> order phase transition
- No known field theory counterpart

# The second example $\rightarrow$ first order phase transition

- As in the previous case there exists minimal temperature  $T_m$
- For the unstable region (red-dashed line) we have  $c_s^2 < 0$



# Holographic spinodal instability



- Modes for  $T \simeq 1.06 T_m$  where  $c_s^2 \simeq -0.1$
- Hydrodynamic mode follows the thermodynamic instability
- Non-hydrodynamic modes have weak momentum dependence



- Thermodynamic instability  $\rightarrow$  dynamical instability
- Converse seems not to be true!  
U. Gursoy, A. Jansen, W. van der Schee, arXiv:1603.07724 [hep-th]
- Non-trivial phase structure limits the applicability of hydrodynamics
- In most cases non-hydro degrees of freedom have very weak dependence on  $k \rightarrow$  „*ultralocality*”
- Extensions to lower couplings and comparison to kinetic theory  
S. Grozdanov, N. Kaplis, A. O. Starinets, arXiv:1605.02173 [hep-th]
- Experimental evidences in cold atoms systems  
J. Brewer, P. Romatschke, Phys. Rev. Lett. **115**, no. 19, 190404 (2015)